Thorn for analysing tilted accretion disks

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Tilted Disks: Motivation and previous work

- Pioneering work in this field by Fragile et al (2005,2006) who have analysed tilted disks in the Cowling (fixed background spacetime) approximation.
- No reason to expect that S_{BH} is aligned with the orbital plane of the NS-BH merger.
- Perform simulations with spacetime evolution to investigate effects of BH tilt to BH+torus evolution.
- Computationally cheaper (due to symmetries) to consider
 S_{BH} and L_{disk} aligned.

Analysis of Twist and Tilt

- We analyse the response of the disk to the tilted BH by two quantities:
- The twist: $\sigma(r) = \angle(\mathbf{S}_{\mathrm{BH}} \times \mathbf{S}_{\mathrm{xy-90}}, \mathrm{P}(\mathbf{J}_{\mathrm{Disk}}(r), \mathbf{S}_{\mathrm{BH}}))$, where

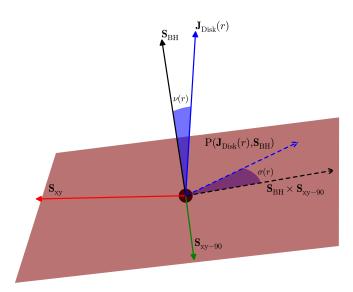
$$P(\mathbf{a}, \mathbf{b}) = \mathbf{a} - \frac{\mathbf{a} \cdot \mathbf{b}}{|\mathbf{b}|^2} \mathbf{b}, \qquad (1)$$

is the projection of vector **a** onto the plane with normal **b**.

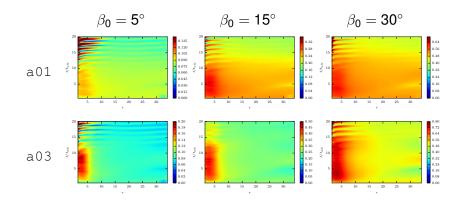
- lacksquare and tilt: $u(r) = \angle(\mathbf{S}_{\mathrm{BH}}, \mathbf{J}_{\mathrm{Disk}}(r))$
- The disk is said to become twisted (warped), if $\sigma(r)$ ($\nu(r)$) vary with r



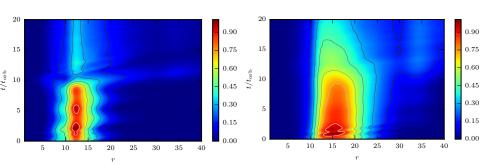
Analysis of Twist and Tilt II



T-r diagrams of complete tilt evolution NC1



Angular momentum transport for C1Ba00 and NC1a00



The End

Thank you for your attention!

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